Development of an interpretive simulation tool for the proton radiography technique

M. C. Levy,^{1,2,a)} D. D. Ryutov,² S. C. Wilks,² J. S. Ross,² C. M. Huntington,² F. Fiuza,² D. A. Martinez,² N. L. Kugland,³ M. G. Baring,⁴ and H.-S. Park²

¹Clarendon Laboratory, University of Oxford, Parks Road, Oxford OX1 3PU, United Kingdom

²Lawrence Livermore National Laboratory, Livermore, California 94551, USA

³Lam Research Corporation, 4400 Cushing Parkway, Fremont, California 94538, USA

⁴Department of Physics and Astronomy, Rice University, Houston, Texas 77005, USA

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Proton radiography is a useful diagnostic of high energy density (HED) plasmas under active theoretical and experimental development. In this paper, we describe a new simulation tool that interacts realistic laser-driven point-like proton sources with three dimensional electromagnetic fields of arbitrary strength and structure and synthesizes the associated high resolution proton radiograph. The present tool's numerical approach captures all relevant physics effects, including effects related to the formation of caustics. Electromagnetic fields can be imported from particle-in-cell or hydrodynamic codes in a streamlined fashion, and a library of electromagnetic field "primitives" is also provided. This latter capability allows users to add a primitive, modify the field strength, rotate a primitive, and so on, while quickly generating a high resolution radiograph at each step. In this way, our tool enables the user to deconstruct features in a radiograph and interpret them in connection to specific underlying electromagnetic field elements. We show an example application of the tool in connection to experimental observations of the Weibel instability in counterstreaming plasmas, using $\sim 10^8$ particles generated from a realistic laser-driven point-like proton source, imaging fields which cover volumes of ~ 10 mm³. Insights derived from this application show that the tool can support understanding of HED plasmas. © 2015 AIP Publishing LLC. [<http://dx.doi.org/10.1063/1.4909536>]

I. INTRODUCTION

Understanding the electromagnetic field generation driven by intense laser-matter interactions is of fundamental importance to high energy density (HED) plasma physics.^{1–3} In this pursuit, the proton radiography diagnostic technique^{4–8} has enjoyed considerable success, providing insight into megagauss-scale electromagnetic fields in inertial confinement fusion (ICF) implosions,^{9–22} large-scale self-organizing electromagnetic field structures in high-velocity counter-streaming plasma flows,²³ magnetic reconnection processes,^{24–26} HED plasma instabilities,^{27–31} and more.

As implemented over the past decade, the proton radiography technique works by passing a low-density point-source-like proton beam through a HED plasma.^{32–40} The proton beam is typically generated using the target normal sheath acceleration (TNSA) process in which an ultraintense short pulse laser ($> 10^{18}$ W cm⁻²) irradiates a solid target, producing a polychromatic proton source with useful energies ranging from ~ 5 –60 MeV.⁴¹ Laser-driven implosions of D³He fusion capsules have also been employed to produce monoenergetic 3 and 14.7 MeV proton sources.^{10,11} The protons generated using either process propagate ballistically from the source to the interaction region containing the HED plasma, deflect from the electromagnetic fields according to the Lorentz force, then travel ballistically to a distant detector where the radiograph, a two dimensional

fluence map, is recorded. Collisional scattering is negligible across a broad range of plasma areal densities 10^{16} – 10^{20} cm⁻². As an example, the stopping power⁴² of a 10 MeV proton beam in carbon is ≈ 41 MeV cm² g⁻¹. Thus, collisional interactions over 1 mm of Carbon having number density 10^{20} cm⁻³ (mass density 2 mg cm⁻³) induce only a 0.1% change in the proton beam energy. Consequently, in these regimes, the fluence map captures the electromagnetic fields alone. Radiography generated in this way is a uniquely high performance diagnostic, imaging HED plasmas with extraordinary spatial resolution of several micrometers and temporal resolution of 1–10 ps.

For the technique's virtues, the general question of how to interpret a radiograph in connection to its underlying electromagnetic fields has remained open. A key challenge stems from the fact that the radiographic image is not a one-to-one electromagnetic field map but rather forms a convolution of the three dimensional fields with the sampling proton properties. Useful aspects of the field geometry have been deduced from qualitative inspection,^{7,8,17,26,43–49} and by means of quantitative estimates based on scalings of the Lorentz force,⁵⁰ when features of the plasma are known.^{13,19,22,24–26,51–55} Recently, analytic theory describing the deconvolution has been developed,⁵⁶ but its application is constrained to simple field geometries and low field strengths, since the general mapping is nonlinear and degenerate.

Numerical simulations can provide insight into a broader range of situations when plasma-dynamical modeling and synthetic radiography modeling tools are used in concert. In the former role, particle-in-cell (PIC) codes are typically

^{a)}Electronic mail: levymc@stanford.edu

FIG. 1. Schematic of key aspects of the proton radiography simulation tool, following the propagation of protons along the z axis from left to right. Parameters controlling the proton source are described in Sec. II B. As an example, specifying the source control vector $\vec{S} = (2, 14.7 \text{ MeV}, 10^9, 1 \text{ cm})$ creates a 14.7 MeV monoenergetic proton source isotropically emitting 1×10^9 protons, imitating a realistic D ^3He source, situated $|z_s| = 1 \text{ cm}$ from the object plane containing the plasma electromagnetic fields. Depicted in the object plane at $z = 0$ is four tilted ellipsoidal magnetic filaments, each having form given by Eq. (3). These fields are created in *PRIME* by specifying the single field control vector $\vec{L} = (4, 2, 2, 100 \mu\text{m}, 50 \mu\text{m}, 500 \mu\text{m}, 0, 0, 1 \text{ MG})$, as described in Sec. II A. A simulated proton radiograph created by the tool is shown in the image plane, situated at $z = z_i$. Details on the field structure underpinning this radiograph are covered in Sec. IV.

employed when kinetic features must be resolved, and hydrodynamic codes when the plasma electron and ion collisional mean-free paths are small relative to the lengthscales of interest. The latter role of simulating the proton radiograph, given the sampling proton properties and the configuration of plasma and electromagnetic field, can be filled using either a “ray trace” or Monte Carlo code.^{8,10,11,15,18,25,33–38,40,57–60} In the ray trace simulation model, a number of straight-line trajectories (rays) are created at the source some distance from the detector, connecting to the detector. The electromagnetic fields along a given ray are path-integrated and a corresponding net Lorentz deflection is applied to that ray’s final position. Ray tracing codes have been widely used not only for protons but also for neutrons, x-rays, and so on, addressing other physical processes such as absorption and scattering. The Monte Carlo, or discretized, numerical approach by contrast represents protons as test particles having the appropriate mass and time-dependent phase space coordinates. As a consequence, all relevant physical processes can be included in the simulation.

In this paper, we describe a new simulation tool that interacts realistic laser-driven point-like proton sources with three dimensional electromagnetic fields of arbitrary strength and structure, using the discretized method, and synthesizes the associated high resolution proton radiograph. The tool, called *PRIME* for Proton Radiograph IMage Exposition, has been developed to support regimes of operation matching those found in the emerging field of HED plasma science. A schematic of the tool’s workings is shown in Fig. 1. The present tool’s implementation of the discretized numerical approach captures all relevant physics effects, including effects related to the formation of caustics.⁵⁶ Electromagnetic fields can be imported from PIC or hydrodynamic codes in a streamlined fashion. A library of electromagnetic field “primitives” is also provided. These primitives can be considered “eigenvectors,” in effect spanning the basis of electromagnetic fields, such that through linear combinations, the user may construct realistic field topologies by hand. This capability allows users to add

a primitive, modify the field strength, rotate a primitive, and so on, while quickly generating a high resolution radiograph at each step. In this way, *PRIME* enables the user to deconstruct features in a radiograph and interpret them in connection to specific underlying electromagnetic field elements. In this paper, we show results from high resolution simulations performed in connection to experimental observations of the Weibel instability in counterstreaming plasmas,²⁸ using $\sim 10^8$ particles generated from a realistic laser-driven point-like proton source, imaging fields which cover volumes of $\sim 10 \text{ mm}^3$. These results show that *PRIME* can support understanding of a broad range of HED plasmas.

II. FEATURES OF *PRIME*

PRIME is a three dimensional simulation tool that we have been developing for modeling HED plasma situations. Both realistic TNSA and D ^3He (14.7 and 3 MeV) laser-driven proton sources have been tested in experimentally realistic configurations and are available to the user. Additionally, the user has the ability to specify a proton source having arbitrary spectral properties. We anticipate that this radiography tool will have two primary uses. The first is in constructing electromagnetic field structures using primitives, guided by the predictions of plasma physics theory and PIC and hydrodynamic simulation results. This approach provides the advantage that fields are free of numerical noise, a key issue arising in kinetic simulations of millimeter and larger-scale plasmas. Here, the user also has the capability to add a primitive, modify the field strength, rotate a primitive, and so on, while quickly generating a high resolution radiograph at each step. In this manner, *PRIME* should provide insights into the crucial question of how to interpret proton radiographs. We also anticipate that synthetic radiographs produced by this tool should become particularly useful in cases where running PIC and hydrodynamic codes is computationally infeasible, and further to guide

these expensive simulations towards larger scales. The second use of this tool will be in quickly and efficiently simulating a high resolution proton radiograph associated with electromagnetic fields exported from PIC and hydrodynamic codes. For this purpose, we have built in the capability to import fields directly from a variety of existing codes (e.g., OSIRIS⁶¹).

Related to the first use, the standard object description in *PRIME* is a three dimensional electromagnetic primitive describing the volumetric field structure. Descriptions, schematics, and simulations of these primitives are given in the Appendix. The user has a number of high level options for inputting these fields, for example, generating a lattice of primitives or programmatically including randomization effects that are enumerated in Sec. II A. By combining primitives together, the user can simulate fields representative of a large number of important HED processes including electrostatic shock waves, magnetized cylindrical shocks, two-stream and other electrostatic instabilities, intense laser-driven $\nabla n \times \nabla T$ “Biermann battery” magnetic fields (for plasma density n and temperature T), magnetic fields creation by collisional current drive in interpenetrating plasma jets, and filamentary magnetic field structures generated via the Weibel instability.^{23,28,62–64}

With respect to numerical schemes, in *PRIME*, we have implemented a modular approach in order to accurately and efficiently simulate the proton radiography technique. This is motivated by the disparate spatial scales characterizing the source–plasma–detector system. The macroscopic volume is vast: the detector typically sweeps out an area $\sim 25 \text{ cm}^2$ and the axial distance between the source and detector, passing through the interaction region containing the HED plasma, can exceed $>10 \text{ cm}$. At the same time, the microscopic field structures associated with the plasma often have spatial scales of $\sim \mu\text{m}$. Simulating the full volume of the cone connecting the source to the detector resolving the electromagnetic fields would require $\sim 10^{14}$ grid cells. This situation clearly exceeds reasonable computational efforts. Therefore to mitigate this issue in *PRIME*, we have divided the system into three regions. The tool covers the source-to-plasma object region, region containing the plasma object itself, and plasma object-to-detector region, as well as the interfaces connecting them. In the plasma region, we are currently using LSP⁶⁵ for the particle push. This provides the additional advantage that scattering models for dense plasmas as well as deflections due to electromagnetic forces can be included. The modular approach in *PRIME* allows a set of electromagnetic fields to be specified, then different proton sources and different detectors to be “hooked up” to these fields in a streamlined manner. For example, in Sec. IV, we show several high resolution proton radiography results of filamentation-instability-driven fields, obtained by keeping the fields unchanged while swapping between realistic proton sources. By allowing users to quickly image the same field configuration using a TNSA proton source, and 3 MeV and 14.7 MeV D ³He proton sources, we show that *PRIME* can help unravel the convolution between the properties of the source and those of the electromagnetic fields. The particle push and other parts of the code have been parallelized in order to take advantage of the Lawrence Livermore National Laboratory (LLNL) Livermore Comput-

ing (LC) Linux architecture, enabling efficient radiography simulations. As such, in order to access the tool at this stage, we ask that interested scientists please correspond with the one of authors.

A. Tools for constructing electromagnetic fields

A robust set of tools is available to the user for constructing electromagnetic fields in *PRIME*. The complete library of analytic electromagnetic field primitives described in Ref. 56 is available to the user, including electrostatic Gaussian ellipsoids, magnetic flux ropes, and magnetostatic Gaussian ellipsoids. Their associated functional forms and schematics are enumerated in the Appendix. While the length scales of the primitives set the grid resolution, the particle push time step is adjusted to the Courant condition⁶⁶ evaluated using the velocity of the protons, enabling efficient and fast simulations. Each primitive is controlled by a set of parameters governing the nominal peak electric (magnetic) field strength E_0 (B_0), the Cartesian position of the primitive’s centroid (x_0, y_0, z_0) with respect to the center of the region containing the HED plasma and two angles θ and ψ indicating the primitive’s polar and azimuthal angles relative to the proton propagation axis $\hat{z}$. The sign of E_0 (B_0) determines whether protons interacting with the primitive will experience a focusing ($E_0, B_0 > 0$) or defocusing force. Spatial extent is specified, taking the ellipsoids, for example, using the parameters a and b representing the major and semi-major axes, respectively. By varying the ratio a/b , the user can produce field structures representative of Weibel instability-driven magnetic filaments, as well as advection laser-driven Biermann battery-like magnetic “pancakes.”^{56,63,64}

The user can construct a field topology featuring an arbitrary number of primitives, each having unique parameters. A number of input methods describing configurations of several primitives are available to the user. At the lowest level, the user specifies a list of N field control vectors each having the form,

$$\vec{G}_n = (\mathcal{P}, x_0, y_0, z_0, \theta, \psi, a, b, E_0(B_0)) \quad (1)$$

which are then transformed by the tool into N volumetric fields in the three dimensional simulation ($n \in [1, N]$). The $\mathcal{P}$ element is an integer mapping to the desired primitive type ($\mathcal{P} = 1$ corresponds to an electrostatic Gaussian ellipsoid, for example; see the Appendix for the complete enumeration). In the simulation, overlapping regions of field have E and B automatically summed.

Higher level input options are also available to the user. To support modeling of periodic systems, a lattice of primitives can be generated programmatically by specifying a single field control vector of the form,

$$\vec{L} = (\mathcal{P}, N_{\text{rows}}, N_{\text{cols}}, d_{\text{rows}}, d_{\text{cols}}, a, b, \theta, \psi, E_0(B_0)). \quad (2)$$

The tool transforms this vector into a body-centered rectangular prism lattice of $N_{\text{rows}} \times N_{\text{cols}}$ primitives of type $\mathcal{P}$. The lattice is centered at the origin of the plasma region and the n th primitive has the centroid position $(x_{0,n}, 0, z_{0,n})$. Rows are oriented along $\hat{z}$ and columns are oriented along $\hat{x}$. N_{rows} is thus the number of primitives in the lattice in $\hat{z}$ and d_{rows} is the

spacing between primitives in $\hat{z}$. Similarly, the “cols” subscript corresponds to periodicity in $\hat{x}$.

To support more realistic field configurations, high level input options that enable randomization effects are also available to the user. By appending the elements $(\delta\theta, \delta\psi, \delta E_0(\delta B_0))$ to the lattice-generating vector $\vec{L}$, the user can programmatically make unique the θ, ψ , and $E_0(B_0)$ values for each primitive. Taking the altitude angle as an example, specifying $\delta\theta = 0$ (or omitting the δ elements) means that $\theta_n = \theta$ for the n th primitive. Randomization effects enter as specifying a nonzero $\delta\theta$ applies the mapping $\theta_n \rightarrow \theta + \delta\theta_n$, where $\delta\theta_n \in [-\delta\theta, \delta\theta]$ is sampled randomly within this interval for each primitive. Individualized parameter effects can be as small or large as desired and are generally quite important since they introduce a realistic asymmetry into the simulation. Indeed, Sec. IV below discusses the significant impact on the resulting proton radiographs of $\delta\theta$ and $\delta\psi$ effects in representing filamentation-instability-driven fields.

B. Specifying source and detector properties

Two methods of proton beam generation are supported, which together offer users the capability to specify sources with arbitrary spectral properties. The first method produces point proton sources. The user can choose a temperature T_p representing TNSA-generated protons having a quasi-Maxwellian distribution. Alternatively, with the first method, users can specify an energy $\mathcal{E}_p$ to generate a monoenergetic point proton source. Specifying $\mathcal{E}_p = 3$ or 14.7 MeV reproduces the properties of protons generated through fusion reactions in intense laser-imploded D ³He capsules. In addition to setting the energy parameter, users also choose the number of protons to simulate N_p and the axial position of the source, $z_s < 0$, relative to the object plane containing the HED plasma at $z_o = 0$. The dimensions of the plasma region ℓ_x, ℓ_y , and ℓ_z (lengths in x, y , and z , respectively) are determined automatically such that they contain the plasma. This region is centered at $(0, 0, 0)$ and is situated between $|x| \leq \ell_x/2, |y| \leq \ell_y/2$ and $|z| \leq \ell_z/2$. The proton source is then instantiated in the simulation at the position $(0, 0, z_s)$ with a phase space distribution corresponding to a point source according to these parameters. In short, a realistic point proton source is created in *PRIME* by specifying a single source control vector of the form $\vec{S} = (S, T_p(\mathcal{E}_p), N_p, z_s)$, where $S = 1$ and the second element is T_p for a TNSA source or $S = 2$ and the second element is $\mathcal{E}_p$ for a monoenergetic source.

In the second proton generation method, the user specifies the proton source “spot” size r_p in addition to N_p and z_s . The source is then instantiated in the simulation at $z = z_s$ with finite transverse size between $x^2 + y^2 \leq r_p^2$. The proton beam divergence and energy distribution are specified through a combination of the beam thermal temperature T_p and a vector drift velocity $\vec{V}_p$. The user can specify spatial variations in both $T_p(x, y)$ and $\vec{V}_p(x, y)$ across the source. This allows a high degree of customization of beam properties, e.g., simulating temperature (Doppler) broadening and finite source size effects in an otherwise monoenergetic point proton source, or reproducing a plane proton source when $T_p = 0$ and $\vec{V}_p(x, y) = (0, 0, \text{const.})$.

To support a range of conditions, the user has the option to specify the detector properties in addition to the source properties. The user may choose the image plane axial position z_i of the detector corresponding to the nominal magnification $M = -z_i/z_s$. The user also can specify the size of the detector and the binning resolution in each transverse direction. If no detector attributes are chosen, the default detector will be instantiated in the simulation with infinite transverse dimensions at $z_i = 10$ cm, with $30 \mu\text{m} \times 30 \mu\text{m}$ resolution in nominal object plane units. By default, the detector records the (x, y) positions of the protons it collects at $z = z_i$, using weighting, that is, irrespective of energy. The user may choose to expand the set of recorded quantities to include the proton velocity $\vec{v} = (v_x, v_y, v_z)$. This capability enables the determination of proton energy deposition within a finite bandwidth, as is useful for many purposes, e.g., for a monoenergetic source, supporting the identification of magnetic $(d/dt \ 1/2m_p \vec{v} \cdot \vec{v} = 0)$ versus electric ($\neq 0$) deflections.

III. BENCHMARKING AGAINST ANALYTIC THEORY

Analytic theory describing the connection between electromagnetic fields and the fluence images produced by sampling protons has been developed in Kugland *et al.*⁵⁶ In this section, predictions of this formalism are compared to results produced by our numerical radiography tool.

Consider a Gaussian ellipsoidal “cocoon” filled with magnetic field having only an azimuthal (φ) component,

$$B_\varphi = B_0 \frac{r}{a} \exp\left(-\frac{r^2}{a^2} - \frac{z^2}{b^2}\right) \quad (3)$$

for radial coordinate r , axial coordinate z , and semi-major and major axes b and a , respectively. For elongated $b > a$ situations, this field structure resembles a single Weibel instability-driven magnetic filament.^{23,63,64} Note that in this representation, B_0 is not a maximum value of the field; the maximum is reached at $r = a/\sqrt{2}$ and is equal to $B_{\text{peak}} = B_0/\sqrt{2}e \approx 0.43B_0$, where e is the natural logarithm base. To create this primitive in the radiography tool, the user specifies the index $\mathcal{P} = 4$ in conjunction with Eq. (1). We assume that the distance from the source to the center of the object is $|z_s| = 1$ cm, the distance from the center to the image plane is $z_i = 10$ cm, proton energy is $\mathcal{E}_p = 1/2m_p v_p^2 = 14.7$ MeV for proton mass m_p and velocity v_p , $a = 100 \mu\text{m}$ and $b = 300 \mu\text{m}$.

This situation is therefore consistent with the paraxial approximation ($a/|z_s| \sim 10^{-2}$). In the analytic evaluation of the proton deflection, we use the smallness of the dimension b compared to the proton gyroradius $\rho \sim 3$ cm for the fields that are needed to form the caustics. This allows us to use a linear approximation: integration of the transverse force over the unperturbed (straight) trajectory within the field structure. The anticipated error of this assumption is less than 10%. With that, we find that deflection angle α is related to the radius r_0 of the point where protons intersect the object plane by

$$\alpha = \mu \frac{r_0}{a} \exp\left(-\frac{r_0^2}{a^2}\right), \quad (4)$$

FIG. 2. Comparison between theory and simulation results in the linear regime. The situations shown here interact a monoenergetic $\mathcal{E}_p = 14.7$ MeV proton source of $|z_s| = 1$ cm, $z_i = 10$ cm with a single magnetic filament of the form given by Eq. (3) having $a = 100$ μm and $b = 300$ μm . (a) illustrates the three dimensional proton-field interaction geometry. The transparent orange surface represents an isocontour of the field magnitude $|B|$, and the colored arrows show the vector B_ϕ field, with both arrow size and color corresponding to field strength. The blue three dimensional arrow indicates the axis of proton propagation. (b) and (c) show simulated radiograph results using $B_0 = 0.2 \times B_{0\text{crit}}$ for the focusing and defocusing cases, respectively, while (e) and (f) use $B_0 = 0.9 \times B_{0\text{crit}}$. The color scale is fixed between images with darker (lighter) regions indicating a surplus (deficit) of protons. (d) and (g) depict normalized lineouts of the proton fluence along $y_0 = 0$ for the $0.2 \times B_{0\text{crit}}$ and $0.9 \times B_{0\text{crit}}$ simulations, respectively. The blue curves correspond to the focusing cases and the yellow curves to the defocusing cases. The simulations agree with the theory predictions of Eqs. (13) and (14), indicated using dashed black curves, to better than 5% in all cases.

where

$$\mu = \frac{\sqrt{\pi}|e|B_0b}{m_p v_p c} \quad (5)$$

is a dimensionless parameter characterizing the interaction and e is the fundamental charge. For the 14.7 MeV proton source $v_p/c = 0.177$ and $\mu = 3.2 \times 10^{-6} B_0[\text{T}] b[\mu\text{m}]$. The position of the point in the image plane is determined by

$$r = z_i \left(-\frac{r_0}{z_s} \mp \alpha(r_0) \right), \quad (6)$$

where the sign “minus” corresponds to the focusing case and the sign “plus” to a defocusing case. The derivative dr/dr_0 is

$$\frac{dr}{dr_0} = -\frac{z_i}{z_s} \left[1 \mp \frac{\mu|z_s|}{a} f(r_0/a) \right], \quad (7)$$

$$f(r_0/a) = \left(1 - 2 \frac{r_0^2}{a^2} \right) e^{-r_0^2/a^2}. \quad (8)$$

For small μ (small magnetic field), the second term is negligible and one has just a uniform magnification. When one increases μ , the condition $dr/dr_0 = 0$ is finally met at some μ_{crit} having different values for the focusing and defocusing cases. For the focusing case, the critical value is

$$\mu_{\text{crit}} = -\frac{a}{z_s}, \quad (9)$$

whereas for the defocusing case,

$$\mu_{\text{crit}} = -\frac{a}{z_s} \frac{e^{3/2}}{2} \approx -2.24 \frac{a}{z_s}. \quad (10)$$

Introducing values of the universal constants, one arrives at the following expressions for the critical magnetic fields:

$$B_{0\text{crit}} [\text{T}] = -8.12 \frac{a}{b} \frac{\sqrt{\mathcal{E}_p[\text{MeV}]}}{z_s[\text{cm}]} \quad (11)$$

and

$$B_{0\text{crit}} [\text{T}] = -18.2 \frac{a}{b} \frac{\sqrt{\mathcal{E}_p[\text{MeV}]}}{z_s[\text{cm}]} \quad (12)$$

for the focusing and defocusing cases, respectively. Using the input parameters for these test cases, we find the fields of 10.38 T and 23.26 T, respectively.

Using Eqs. (4)–(8), the intensity distribution in the image plane for μ smaller than critical can be presented in parametric form as

$$\frac{I}{I_0} = \left| e^{-2r^2} \left(\nu \mp e^{t^2} \right) \left(e^{t^2} \mp \nu(1 - 2t^2) \right) \right|^{-1}, \quad (13)$$

$$\frac{r}{R} = t \left| 1 \mp \nu e^{-t^2} \right|, \quad \nu \equiv -\frac{\mu z_s}{a}, \quad (14)$$

for parameter t . Here, I_0 is the intensity in the center of the image plane in the absence of an object and $R = -z_i a / z_s$.

One can also plot intensity distributions for the fields exceeding critical values. In order to do so, the amplitude-limiting factor ϵ as described in Ref. 56 must be accounted for in Eq. (13). The appropriate parametric relation for the normalized image plane intensity is then given by

$$\frac{I}{I_0} = \frac{1 + \epsilon}{\epsilon + \left| e^{-2r^2} \left(\nu \mp e^{t^2} \right) \left(e^{t^2} \mp \nu(1 - 2t^2) \right) \right|} \quad (15)$$

in concert with Eq. (14) for r/R . The transformation $z_s \rightarrow \tilde{z}_s$, where $\tilde{z}_s = z_s z_i / (z_i - z_s)$ enhances the accuracy of Eqs. (13)–(15) by relaxing constraints on the relationship between z_i and z_s .⁵⁶

We now validate the synthetic radiographs produced by the numerical tool through comparison to Eqs. (13)–(15). Fig. 2 shows the results of this procedure for four simulations in the linear regime. (b) and (c) show synthetic proton radiographs generated by the tool using $B_0 = 0.2 \times B_{0\text{crit}}$ for the focusing and defocusing cases, respectively. The color scale is fixed between images (and Figs. 2–6), with darker (lighter) regions indicating a surplus (deficit) of protons. The spatial

FIG. 3. Comparison between theory and simulation results in the caustic regime. In these simulations, a monoenergetic $\mathcal{E}_p = 14.7$ MeV proton source of $|z_s| = 1$ cm, $z_t = 10$ cm interacts with a single magnetic filament of the form given by Eq. (3) having $a = 100$ μm and $b = 300$ μm . The interaction geometry is identical to that shown in Fig. 2(a). (a) and (b) show simulated radiograph results using $B_0 = 2 \times B_{0\text{crit}}$ for the focusing and defocusing cases, respectively. The color scale is fixed between images with darker (lighter) regions indicating a surplus (deficit) of protons. (d) shows normalized lineouts of the simulated proton fluence along $y_0 = 0$ with the blue curve corresponding to the focusing case and the yellow curve corresponding to the defocusing case. The complete analytic results formed by summing over all three branches of each curve in (c) are indicated by the dashed black lines in (d), exhibiting close agreement with the simulation results. (c) shows the multi-branched caustic structures predicted by parametric Eqs. (14) and (15) using $\epsilon = 0.5$.

coordinates are provided in nominal object plane units x_0 and y_0 , i.e., $1/M \times x, y$. (d) depicts normalized lineouts of the proton fluence along $y_0 = 0$ with the blue curve corresponding to the focusing simulation and the yellow curve corresponding to the defocusing simulation. The black dashed curves correspond to analytic theory from Eq. (14). (e)-(g) show the same set of plots for simulations and theory corresponding to the field strength $B_0 = 0.9 \times B_{0\text{crit}}$. Panels (d) and (g) highlight the excellent agreement between theory and the simulated radiographs across conditions.

Fig. 3 shows results comparing simulations to the predictions of Eqs. (14) and (15) for proton imaging in the nonlinear regime. (a) and (b) show the synthetic proton radiographs having nonlinear field strength $B_0 = 2 \times B_{0\text{crit}}$ for the focusing and defocusing cases, respectively. (c) shows the multi-branched caustic structures predicted by parametric equations (14) and (15). (d) shows normalized lineouts of the simulated proton fluence along $y_0 = 0$ with the blue curve corresponding to the focusing case and the yellow curve corresponding to the defocusing case. The complete analytic results formed by summing over all three branches of each curve in (c) are indicated by the dashed black lines in (d). Plots (c) and (d) use $\epsilon = 0.5$, a value chosen so that the magnitudes of the I/I_0 analytical curves most closely match the simulation data. This is necessary in this situation since, using a point proton source, for $\epsilon = 0$, analytically the caustic intensities tend towards infinity.⁵⁶ Recent germane experimental results have suggested that $I/I_0 \sim 3$, in practice,²⁸ illustrating the importance of $\epsilon > 0$ accounting for finite resolution effects. Consistent with this finding, (d) shows that the simulation output closely matches the analytics, bolstering confidence in its numerical fidelity.

IV. APPLICATION TO THE FILAMENTATION INSTABILITY IN MILLIMETER-SCALE HED PLASMAS

We have developed *PRIME* in connection to laboratory astrophysics experiments performed by the ACSEL collaboration.^{23,67} These experiments use powerful lasers to create high velocity plasmas flows by ablating the surface of plastic (CH_2) targets. In a typical experiment, two such targets are setup opposing one another and illuminated with laser light to study properties of the colliding plasma plumes. For our purposes, here, the typical plasma parameters^{56,67-69} are $n_e = 1 \times 10^{19} \text{ cm}^{-3}$, $T_e = T_i = 1$ keV, $v_{\text{flow}} = 8 \times 10^7$ cm/s. In the interaction between the two flows, it is believed that the Weibel filamentation instability⁷⁰ plays an important role. Indeed, Weibel-like filamentary structures appearing in proton radiographs of the interaction have recently been reported.^{28,29} Yet for the reasons described above, the challenge to discern the fields from their radiograph, i.e., to determine the extent to which filamentary magnetic fields produce filamentary radiograph structures, persists. Realistic situations introduce further questions: will protons traversing the hundreds of magnetic filaments expected in a realistic situation produce a coherent radiograph, or will they scatter; how important are density and temperature heterogeneities expected in the plasma flows; what is role of field strength as the filaments grow over time; and ultimately if a coherent radiograph can be produced how does its periodicity correspond to that of the underlying fields. Resolving these complications will evidently require many simulations, and due to the plasma's $\sim 10 \text{ mm}^3$ scale computational expense implies that multidimensional hydrodynamic and PIC simulations will not be ideally suited to this purpose.

FIG. 4. Comparison of simulation fields to the relevant transverse Weibel instability modes. (a) shows the object plane configuration space of the magnetic filament centroids (at $y_0 = 0$) for three simulations: sim. 18 (blue), sim. 45 (orange), and sim. 99 (green), as described in the text. (b) compares the k -space of these filament centroids across $\hat{x}$ at $z_0 = 0$ to the germane theoretical instability growth rates. The colored curves correspond to normalized Fourier transformations of the natural logarithm $\log B_x(k)$ from (a). The black curves correspond to the normalized instability growth rates $\Gamma(k)$ for collisionless Carbon (dotted), collisional Carbon (dashed), and collisional CH_2 flows (solid) from Eqs. (16) and (17).

Our purpose here is to show that, using electromagnetic primitives to construct representative filamentary fields, *PRIME* simulations can provide insight into this situation. To this end, we address a subset of these questions in this section.

We construct a representative field topology, guided by the reported experimental conditions,²⁸ using many dozens of magnetostatic Gaussian “cocoon” of the form given in Eq. (3). The experimental results imply that filaments form within a ~ 1.5 mm radius cylinder in the interaction midplane, with axial coordinate directed between the opposing plastic targets. We model this as a “forest” of 260 filaments each instantiated with a random centroid position in the $x_0 - z_0$ plane (at $y_0 = 0$, oriented along $\hat{y}$) within $x_0^2 + z_0^2 \leq (1.5 \text{ mm})^2$. Experimental conditions also imply that $c/\omega_{pi} = 100 \text{ } \mu\text{m}$ and the axial length of the cylinder containing the filaments ~ 0.5 mm, so in the simulation each filament has $a = 50 \text{ } \mu\text{m}$, $b = 500 \text{ } \mu\text{m}$, meaning that the inverse wavenumber of the filament centroids is nominally $2a = c/\omega_{pi}$. We further use randomized tilt parameters $\delta\theta, \delta\psi = 15^\circ$ to account for natural density perturbations occurring in the plasma. Since these perturbations affecting the filament growth can be expected to vary between experiments, and since we are interested in determining whether filamentary structures in the radiographs are a robust signature of filamentary magnetic fields, we instantiate this setup in three distinct simulations. That is, we perform three simulations pursuant to these conditions, meaning that the filament centroid positions in $x_0 - z_0$ and

the individualized tilts of filaments will vary between simulations, while each filament a and b and the density of filaments across simulations are constant. The three simulation geometries are shown in Fig. 4(a). This plot shows the positioning of filament centroids, which varies between simulations in a randomized fashion, as well as the high density of filaments, which is held constant at $\sim 75 \text{ mm}^{-3}$ across simulations.

To see that these simulation conditions form a reasonable approximation of experimental conditions, it is instructive to consider the relevant Weibel instability growth rates. For the purely transverse mode, the collisionless dispersion relation is given by

$$k^2 + \frac{\Gamma\sigma}{|k|U_e + \Gamma} + \frac{\Gamma}{|k|U_i + \Gamma} = \frac{k^2}{\Gamma^2 + \frac{3k^2S}{5}}, \quad (16)$$

where Γ is the growth rate normalized to $v_{\text{flow}} \omega_{pi}/c$, k is the wave number normalized to ω_{pi}/c , $U_{e,i} = v_{T e,i}/(\sqrt{\pi}v_{\text{flow}})$ for thermal velocity v_T , $S = 0.014 T_i [\text{keV}]$, and $\sigma = Am_p/(Zm_e)$ for atomic mass A and charge state Z .^{69,71} The dispersion relation accounting for inter-flow collisional effects⁶⁹ can be formulated as

$$k^2 + \frac{\Gamma}{\Gamma + k^2V_s} + \frac{\sigma(\Gamma + k^2R)}{\Gamma + k^2V_{se}} = \frac{k^2}{\Gamma(k^2V_b + \Gamma) + k^2S}, \quad (17)$$

where $R = 0.00106/(T_e [\text{keV}])^{3/2}$, $V_s = 0.0175 (T_i [\text{keV}])^{5/2}$, $V_b = 0.0253 (T_i [\text{keV}])^{5/2}$, $V_{se} = 64 (T_e [\text{keV}])^{5/2}$.

Equations (16) and (17) provide physical references for filament periodicity in the simulations. Fig. 4(b) shows the simulation fields in relation to the normalized Γ curves for collisionless Carbon flows (dotted black), collisional Carbon flows (dashed black), and collisional CH_2 flows (solid black) in which the light ions exhibit a stabilizing effect on the instability growth. The Γ calculations assume plasma states of full ionization consistent with typical conditions that $T_e = T_i = 1 \text{ keV}$, $v_{\text{flow}} = 8 \times 10^7 \text{ cm/s}$, and their depictions in Fig. 4(b) indicate the transverse Weibel modes which can be expected to grow most rapidly in the plasma. Since $\hat{z}$ is the axis of proton propagation, the protons will deflect most strongly from the filamentary B_x fields. The colored curves correspond to normalized Fourier transformations $\log B_x(k)$ across $\hat{x}$ at the simulation midplane for the three simulations: sim. 18 (blue), sim. 45 (orange), and sim. 99 (green). Here, $\log$ denotes the natural logarithm and B_x forms the representative field quantity, since the filament orientation along $\hat{y}$ means that B_y reflects only $\delta\theta, \delta\psi$ effects. From Fig. 4(b), it is clear that the simulations provide an imperfect but reasonable approximation of the k -vectors which can be expected in the experimental situation.

Having described the simulation setup, we now analyze the synthetic proton radiographs generated by *PRIME* for these cases. First, we consider the roles of the magnetic field strength and proton beam energy for a single field configuration. Fig. 5 shows the simulated proton radiographs for two values of B_0 and two values of $\mathcal{E}_p$. (a) corresponding to the $B_0 = 1 \text{ MG}$ ($B_{\text{peak}} \approx 0.4 \text{ MG}$) field strength and $\mathcal{E}_p$

FIG. 5. Effects of field strength B_0 and probing proton energy $\mathcal{E}_p$ on the simulated proton radiograph results. All panels here correspond to the sim. 99 configuration indicated by the green curve in Fig. 4. (a) and (b) correspond to probing proton energies of $\mathcal{E}_p = 14.7$ MeV and (c) and (d) to $\mathcal{E}_p = 3$ MeV. (a) and (c) correspond to $B_0 = 1$ MG (meaning a peak simulation field of 0.4 MG), and (b) and (d) correspond to $B_0 = 0.3$ MG.

= 14.7 MeV proton source closely approximates the calculated field values and the experimental conditions reported on in Ref. 28. In this simulated radiograph, we observe coherent, predominantly vertical filamentary features striated along the plasma flow axis ($\hat{y}$). This fact is striking since according to Ref. 56, protons should deflect in a nonlinear fashion from each of several dozen filamentary field structures on their path to the detector. Through examination of (b)-(d), it is clear that these filamentary features persist across a variety of configurations. Comparison of (a) and (b) further shows that a reduction in field strength causes an apparent contraction of the plasma flow interaction region. The potential conflation in this regard forms an important consideration for experimental diagnosis. We also note that the tilting of the field filaments, a feature expected in realistic situations, plays an important role in the simulated radiograph signal. In additional simulations not presented here, we observed that $\delta\psi, \delta\theta \rightarrow 0$ reduces the fluence amplitude of structures present in the radiograph by a factor of three or more.

To examine the robustness of filamentary radiograph structures, we examine the sim. 18 and sim. 45 field configurations. Fig. 6 depicts these images, which are seen to clearly exhibit similar coherent, predominantly vertical filamentary features. In order to characterize the relationship between the field periodicity and the radiograph periodicity, we have analyzed lineouts of the proton fluence along $y_0 = 0$ for each of the simulations. Fig. 7(a) shows the magnitude of the Fourier-transformed periodicity from each radiograph. In (b)-(d), these radiograph periodicities (solid lines) are compared to the underlying magnetic field periodicities (dashed lines). From these figures, it is clear that the radiograph signal is shifted to much shorter wavelengths than those found in the simulation. Furthermore, the radiograph signal is negligible at the low k -values which dominate the magnetic field spectra. These results show that, at minimum for the cases considered here, filamentary structures in proton radiographs are a qualitative signature of Weibel instability-like filamentary magnetic fields. Future work will focus on parsing the quan-

FIG. 6. Synthetic proton radiographs for (a) sim. 18 and (b) sim. 45. Across simulations $B_0 = 1$ MG and $\mathcal{E}_p = 14.7$ MeV. The radiograph corresponding to these conditions for sim. 99 is shown in Fig. 5(a).

FIG. 7. Comparison between field periodicity and proton radiograph image periodicity. Solid lines correspond to Fourier-transformed lineouts at $y_0=0$ of the synthetic radiographs shown in Fig. 6. The dashed curves in (b)-(d) correspond to $2\pi/k$ for the simulation k -vectors shown in Fig. 4.

titative relationship between the field and radiograph periodicities, a task which exceeds the illustrative scope of this section.

V. CONCLUSIONS

We have presented a new simulation tool for interpreting proton radiography of HED plasmas. The present tool's numerical approach captures all relevant physics effects, including effects related to the formation of caustics. Electromagnetic fields can be imported from PIC or hydrodynamic codes in a streamlined fashion. A library of electromagnetic field “primitives” is also provided. These primitives can be considered “eigenvectors,” in effect spanning the basis of electromagnetic fields, such that through linear combinations, the user may construct realistic field topologies by hand. This

capability allows users to add a primitive, modify the field strength, rotate a primitive, and so on, while quickly generating a high resolution radiograph at each step. In this way, *PRIME* enables the user to deconstruct features in a radiograph and interpret them in connection to specific underlying electromagnetic field elements. We have applied the tool in connection to experimental observations of the Weibel instability in counterstreaming plasmas, using $\sim 10^8$ particles generated from a realistic laser-driven point-like proton source, imaging fields which cover volumes of $\sim 10 \text{ mm}^3$. Insights derived from this application indicate that tilting of magnetic filaments plays a significant role in setting the proton image; field strength tends to affect the apparent axial lengthscale over which the filamentation instability is active; and coherent imaging is possible in the sense that filamentary structures are observed in radiographs as a signature of the Weibel fields, at least for

TABLE I. Basic descriptions of the electromagnetic field primitives available in *PRIME*.

$\mathcal{P}$	Electric potential or magnetic field vector	Description and comments
1	$\frac{\phi}{\phi_0} = \exp\left(-\frac{x_0^2 + y_0^2}{a^2} - \frac{z_0^2}{b^2}\right)$	Electrostatic Gaussian ellipsoid; representative of electronic density variations in collisional plasmas.
2	$\frac{\phi}{\phi_0} = \frac{1}{2} e^{-\frac{x_0^2 + z_0^2}{a^2}} \left(1 - \operatorname{erf}\left(\frac{y_0}{d}\right)\right) \log\left(\frac{n}{n_0}\right)$	Electrostatic quasi-planar shock propagating along a cylinder in $\hat{y}$; describes shocks having thickness $d \gg \lambda_D$ (λ_D being the Debye length), i.e., electric fields are ambipolar in origin; plasma density decreases along the radial coordinate; and n/n_0 is the ratio of post-shock to pre-shock electron densities.
3	$\frac{B_y}{B_0} = \exp\left(-\frac{x_0^2 + z_0^2}{a^2}\right)$	Magnetic flux rope; characteristic of objects appearing in laser-generated plasma jets.
4	$\frac{\vec{B}}{B_0} = \frac{1}{a} \exp\left(-\frac{x_0^2 + y_0^2}{a^2} - \frac{z_0^2}{b^2}\right) (-y_0, x_0, 0)$	Magnetostatic Gaussian ellipsoid; for $a/b \gtrsim 1$, characteristic of mechanisms of magnetic field generation such as collisional current drive in interpenetrating plasma jets and the “Biermann battery.” ^{62,63} For $a/b < 1$, characteristic of a current filament surrounded by a return current, relevant to Weibel instability-driven magnetic fields (see Secs. III and IV).

TABLE II. Schematics and simulated radiographs of the available electromagnetic field primitives for a variety of conditions.

$\mathcal{P}$	Object plane field schematic and comments	Simulated radiograph
1	<p>$a/b=1; E_0 = -2000 \text{ kV cm}^{-1}$</p>
1	<p>$a/b=1/3; E_0 = -200 \text{ kV cm}^{-1}$</p>
2	<p>$a = 300 \mu\text{m}; n/n_0 = 4; d = 10 \mu\text{m};$ $E_0 = -200 \text{ kV cm}^{-1}$</p>

TABLE II. (Continued.)

$\mathcal{P}$	Object plane field schematic and comments	Simulated radiograph
3	<p>$a = 100\mu\text{m}; B_0 = 5\text{ T}$</p>
4	<p>$a/b=1; B_0 = 5\text{ T}$</p>
4	<p>$a/b=1/3; B_0 = 5\text{ T}$</p>

TABLE II. (Continued.)

$\mathcal{P}$	Object plane field schematic and comments	Simulated radiograph
4	<p>$a/b=3; B_0 = 10\text{ T}$</p>

the cases considered here. These results show that *PRIME* can support understanding of HED plasmas.

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APPENDIX: DESCRIPTIONS, SCHEMATICS, AND SIMULATIONS OF ELECTROMAGNETIC FIELD PRIMITIVES

In order to develop intuition connecting a proton radiograph image to its underlying electromagnetic field primitive, in this appendix, we enumerate the set of available primitives and provide representative schematics and simulations. Table I represents the four basic field primitives, showing the physical descriptions, unrotated functional forms of the electric potential ϕ or magnetic field vector $\vec{B}$, as well as the corresponding $\mathcal{P}$ index used to invoke each primitive in the simulation. The electric field $\vec{E} = -\nabla\phi$ is obtained in the standard way, and rotation of each primitive along two axes is enabled by specifying the θ and ψ elements of the field control vectors given by Eqs. (1) and (2). We maintain the coordinate systems and notations described in Secs. I–V, so that x_0, y_0, z_0 denote object plane coordinates, ϕ_0 (B_0) is the nominal peak electric potential (magnetic field) of each primitive, and a and b are the major and semi-major axes of each primitive, respectively. For the second row of Table I, the primitive represents a quasi-planar shock propagating along a cylinder in $\hat{y}$ with plasma density decreasing along the radial coordinate, where erf is the Gaussian error function, $\log$ is the natural logarithm, and the

shock thickness $d = 2\sqrt{\log 2}F$ for full width at half maximum of the shock potential F .⁵⁶

The descriptive information presented in Table I is complemented by each primitive's visual representation in the object and image planes. This information is shown in Table II for a variety of conditions relevant to HED plasmas. In this latter table, the primitives' index $\mathcal{P}$ is shown in the first column, object plane schematic in the second column, and (image plane) simulated proton radiograph in the third column. In each schematic, the transparent orange surface represents an isocontour of the field magnitude, and the colored arrows show the vector field, with both arrow size and color corresponding to field strength. The first row's schematic highlights the geometry of the three dimensional proton-field interaction, with the blue three dimensional arrow at the lower z boundary indicating the direction of proton propagation (along $\hat{z}$). This geometry is maintained for all schematics and simulations shown in the table. Supporting clarity of interpretation, in each simulation, we use an identical point source of monoenergetic 14.7 MeV protons, isotropically emitting 1×10^9 particles, imitating a realistic D ^3He source situated $|z_s| = 1\text{ cm}$ from the object plane. In *PRIME*, this source is instantiated in the simulation using the source control vector $\vec{S} = (2, 14.7\text{ MeV}, 10^9, 1\text{ cm})$, as is covered in Sec. II B. As an example of field instantiation, the electric field shown in the second column of the first row is created in the simulation by specifying the field control vector according to Eq. (1) of $\vec{G}_1 = (1, 0, 0, 0, 0, 0, 100\text{ }\mu\text{m}, 100\text{ }\mu\text{m}, -200\text{ kV cm}^{-1})$, as is covered in Sec. II A. For each case, the simulated proton radiograph is shown using nominal object plane units, using the default detector having a magnification factor of ten. The fiducial peak field values for each case are labeled in the second column of the table. The characteristic strengths with which each primitive deflects the protons sampling it is highlighted by the scales of the radiograph fluence shown in the third column. Furthermore, the set of characteristic proton radiographs produced by imaging these primitives is expanded by adjusting the θ, ψ elements of $\vec{G}_n$.⁵⁶ Through tuning of the elements of the $\vec{G}_n$ field control vectors, the primitives enumerated

in this section effectively span the basis of electromagnetic fields, such that by $\sum_n \vec{G}_n$, the user may construct realistic field topologies by hand.

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